The DarkSide WIMP Search Program





Ben Loer 8th Patras Workshop on Axions, WIMPs and WISPs July 18, 2012



Outline



- Argon as a dark matter target
- The DarkSide dual phase argon TPC design
- Low radioactivity argon
- Recent results from the DarkSide-10 prototype
- DarkSide-50 Schedule and outlook



The DarkSide Collaboration



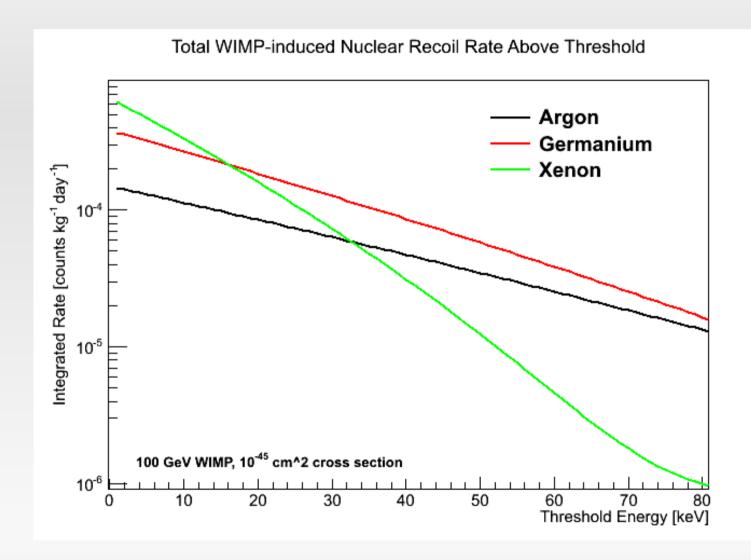
D. Akimov^k, T. Alexander^d, D. Alton^a, K. Arisaka^v, H.O. Back^m, P. Beltrame^v, J. Benziger^l, A. Bolozdynya^k G. Bonfinii, A. Brigattir, J. Brodskym, L. Cadonatix, F. Calapricem, A. Candelai, H. Caom, P. Cavalcantei, A. Chavarria^m, A. Chepurnov^j, S. Chidzik^m, D. Cline^v, A.G. Cocco^s, C. Condon^m, D. DAngelo^r, S. Davini^w, E. De Haas^m, A. Derbinⁿ, G. Di Pietro^r, I. Dratchnevⁿ, D. Durben^b, A. Empl^w, A. Etenko^k, A. Fan^v, G. Fiorillo^s K. Fomenkoⁱ, G. Forster^x, F. Gabriele^m, C. Galbiati^m, S. Gazzanaⁱ, C. Ghag^p, C. Ghianoⁱ, A. Goretti^m, L. Grandi^m M. Gromov¹, M. Guan^e, C. Guo^e, G. Guray^m, E. V. Hungerford^w, Al. Ianniⁱ, An. Ianni^m, A. Kayunovⁿ, K. Keeter^b, C. Kendziora^d, S. Kidner^y, R. Klemmer^m, V. Kobychev^f, G. Koh^m, M. Komor^m, D. Korablev^h, G. Korga^w, E. Shields^m, P. Lie, B. Loerd, P. Lombardir, C. Loveo, L. Ludhovar, L. Lukyanchenko¹, A. Lund^x, K. Lung^y, Y. Mae, I. Machulin^k, J. Maricic^c, C.J. Martoff^o, Y. Meng^v, E. Meroni^r, P. Meyers^m, T. Mohayai^m, D. Montanari^d M. Montuschiⁱ, P. Mosteiro^m, B. Mount^b, V. Muratovaⁿ, A. Nelson^m, A. Nemtzow^x, N. Nurakhov^k, M. Orsiniⁱ F. Ortica^t, M. Pallavicini^q, E. Pantic^v, S. Parmeggiano^r, R. Parsells^m, N. Pelliccia^t, L. Perasso^q, F. Perfetto^s, L. Pinsky^w, A. Pocar^x, S. Pordes^d, K. Randle^x, G. Ranucci^r, A. Razetoⁱ, A. Romani^t, B. Rossi^{m,s}, N. Rossi^{i,m} S.D. Rountreey, P. Saggesei, R. Saldanhai, C. Salvoq, W. Sandsm, M. Seigaru, D. Semenovn, M. Skorokhvatovk, O. Smirnovh, A. Sotnikovh, S. Sukhotink, Y. Suvarov, R. Tartaglia, J. Tatarowiczo, G. Testera, A. Teymourian, J. Thompson^b, E. Unzhakovⁿ, R.B. Vogelaar^y, H. Wang^v, S. Westerdale^m, M. Wojcik^g, A. Wright^m, J. Xu^m C. Yange, S. Zavatarelliq, M. Zehfusb, W. Zhonge, G. Zuzelg

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Why argon?





Argon as a Dark Matter target

- Why an argon TPC?
 - Very efficient scintillator (40 photons/keVee, long absorption length, can be purified easily)
 - Only moderate cryogenic requirements (87 K)
 - Pulse shape and scintillation/ionization ratio give 2 independent background rejection handles
 - TPC gives accurate 3D position reconstruction to reject surface events



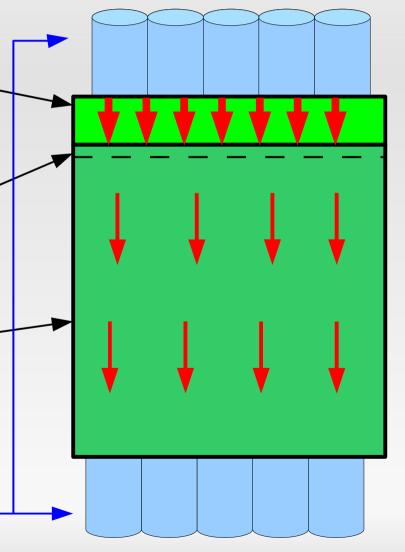


Gaseous argon
 volume with
 multiplication field
 ~3-6 kV/cm

 Grid plane allows for discontinuity in field

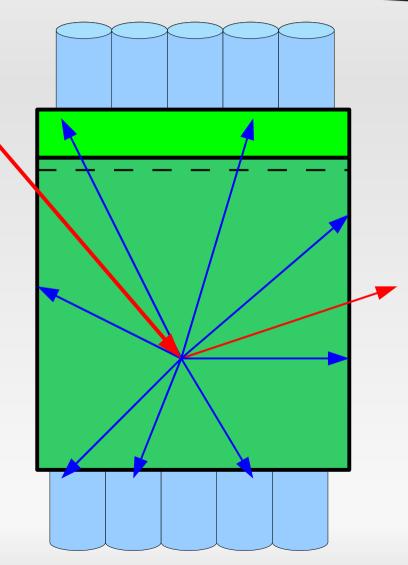
 Liquid argon active volume with drift field ~1 kV/cm

Photodetectors





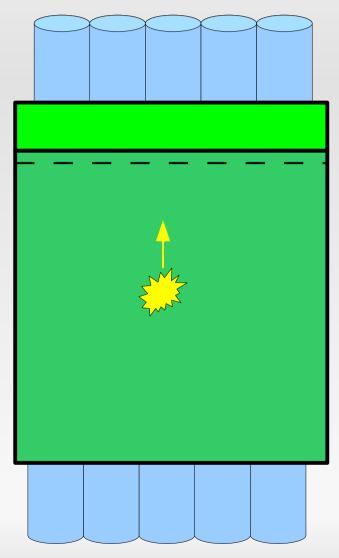
- An energetic particle (WIMP, gamma, etc) interacts in the liquid volume
- Prompt scintillation light (S1) is emitted from the point of interaction
- Gives time and energy of the event







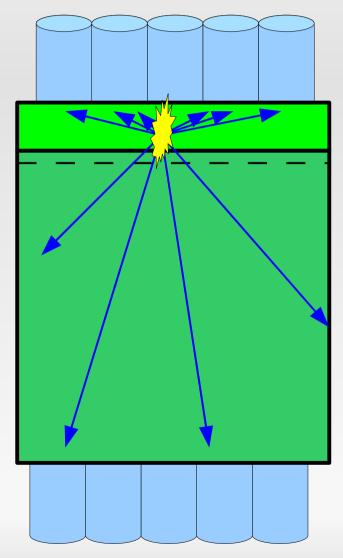
- Interaction also produces free electrons from ionization
- Electric field causes them to drift upward at ~1mm/microsec







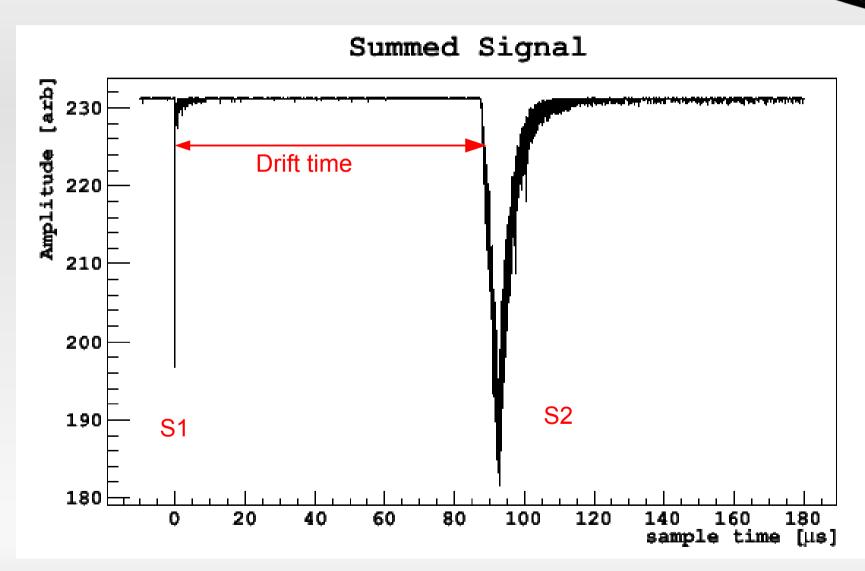
- High electric field in the gas region gives electrons energy to excite secondary scintillation (S2)
- Size of S2 gives amount of ionization
- Light collection ratio in PMTs gives x-y position
- Time between S1 and S2 gives z position





Typical Event

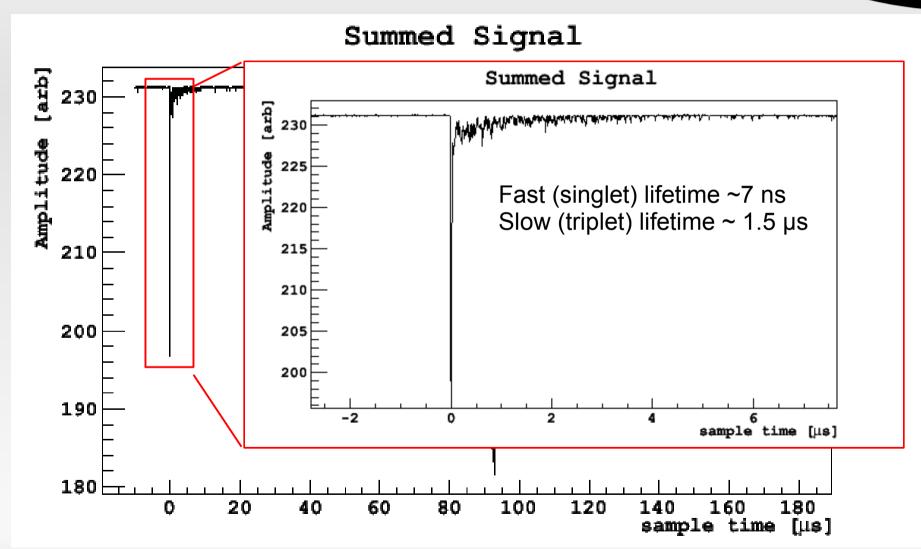






Typical Event



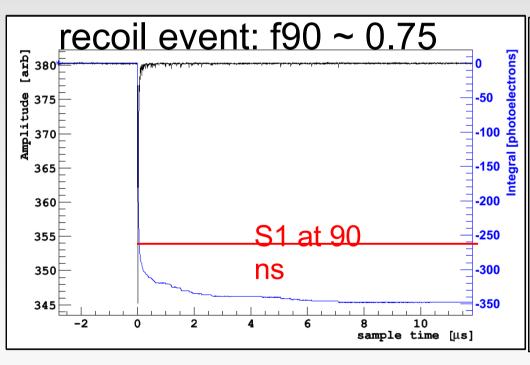


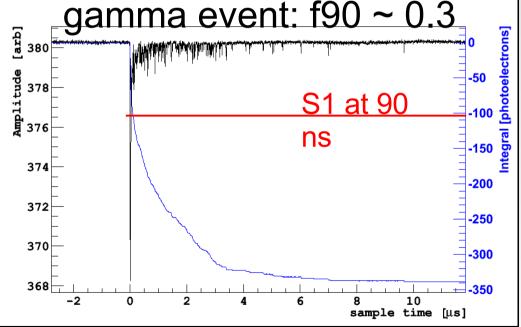


Pulse Shape Discrimination



 f90: the fraction of S1 light that arrives within the first 90 ns

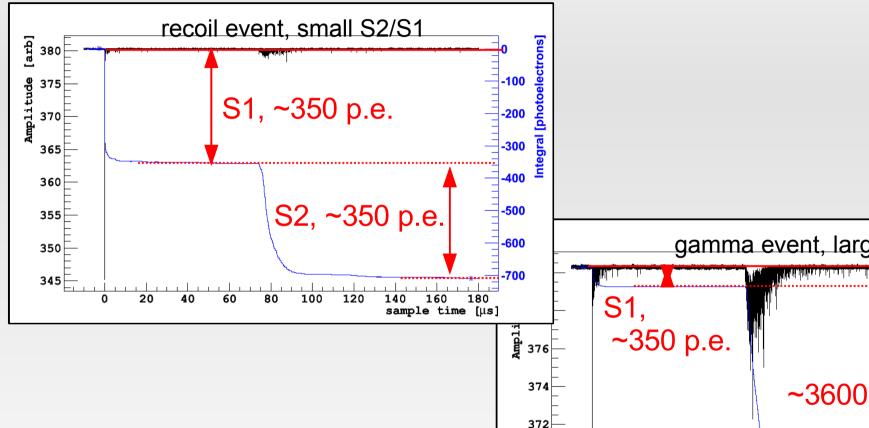


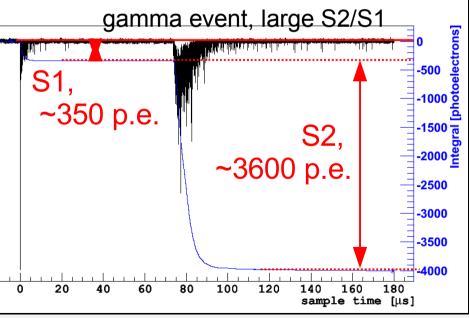




Scintillation/Ionization Ratio



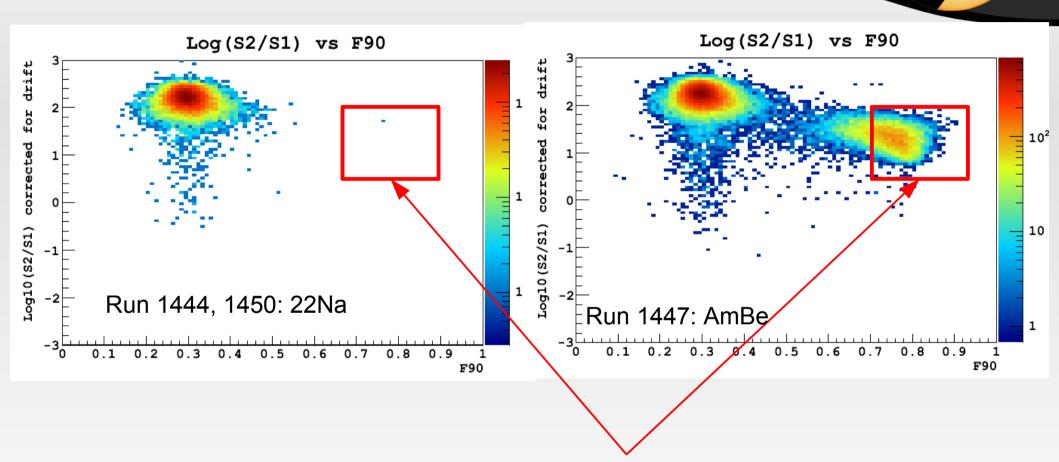






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DarkSide-10 Analysis: Combined Background Rejection



Example nuclear recoil acceptance bounds



DarkSide-10



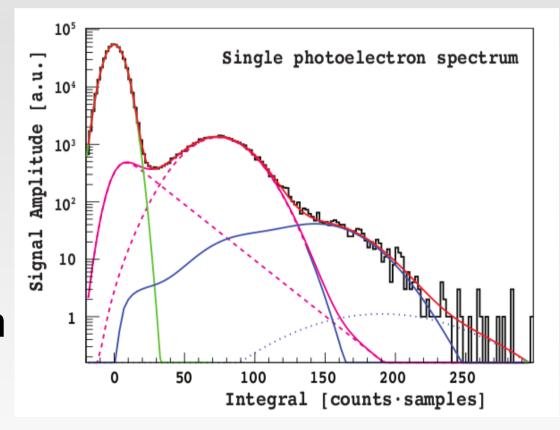




Photomultiplier Calibration



- New to market Hamamatsu 11065 PMTs
- Cryogenic capable
- High quantum efficiency (~35%!)
- Single p.e. spectrum seems to contain exponential component



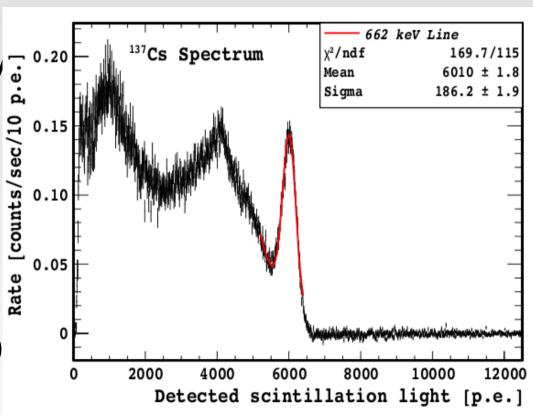
arxiv:1204.6218



DarkSide-10 Light yield



- Measured ~9 p.e./keV light yield, ~4% resolution at 662 keVa
- Best reported from liquid argon detector
- More light means lower energy threshold, better PSD



arxiv:1204.6218



Underground argon

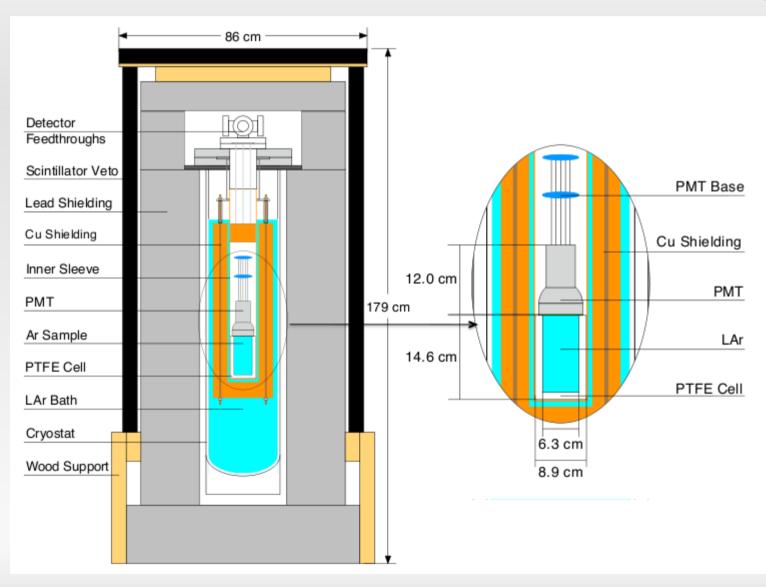


- Atmoshperic argon contains Ar-39, which undergoes beta decay with Q=565 keV, half-life 269 y, at ~1 Bq/kg natural argon
- TPCs limited at ton scale by pileup
- Ar-39 is produced by cosmic rays in the upper atmosphere, so underground argon might have reduced Ar-39 content
- But it can be produced by n,p reactions on K-40, so different sites must be tested



Underground argon measurement



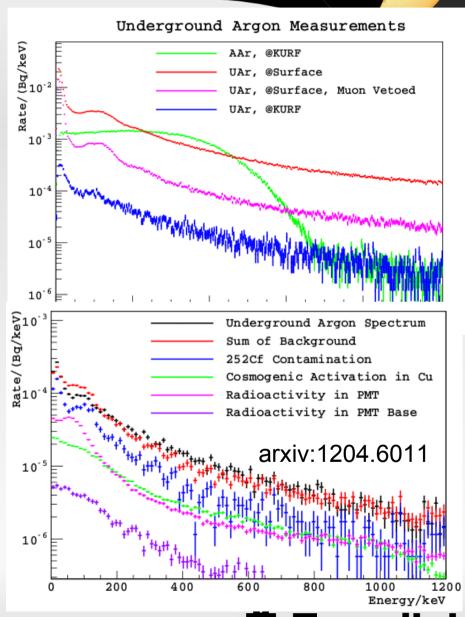




Underground argon decay rate



- Detector operated at Kimbalton Underground Research Facility (KURF)
- Ar-39 rate < 0.65% of atmospheric rate
- Suitable for multi tonscale TPCs



Underground argon extraction



- Stage 1 is pressure swing adsorption
- Boosts few hundred ppm of argon in stream to few percent
- Producing ~0.5 kg/day,
- ~100 kg produced so far



arxiv:1204.6024



Underground argon extraction



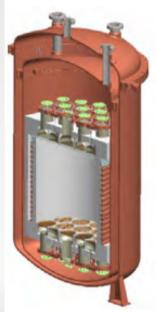
- Stage 2 is a cryogenic distillation column
- Purifies PSA output to <0.05% impurities at ~1 kg/day (limit of RGA, not distillation)
- Plans to upgrade both stages to produce 50 kg/day

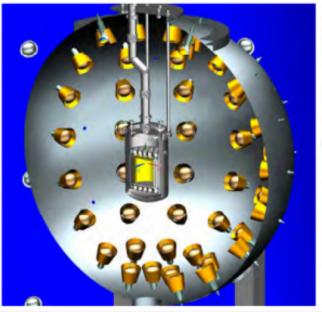


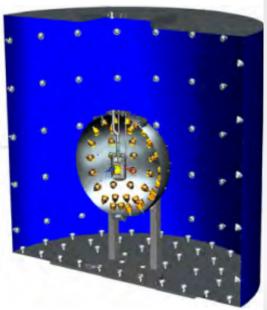


DarkSide-50: Coming soon to LNGS

- DARKSIDE
- 50 kg underground argon (~33kg fiducial)
- 4m diameter boron-loaded liquid scintillator neutron veto (estimated >99.5% veto efficiency for radiogenic neutrons)
- 11m diameter water shield/muon veto



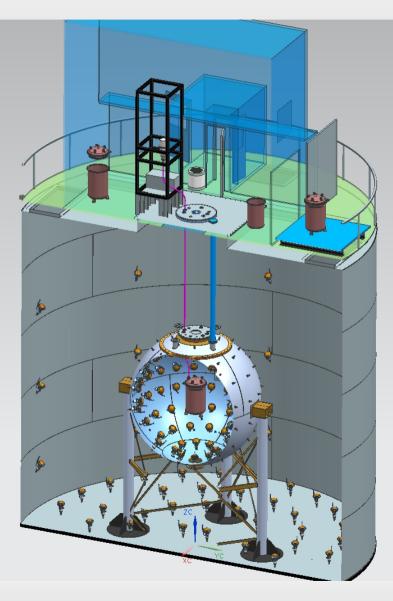






DarkSide-50: Coming soon to LNGS









DarkSide-50 Schedule



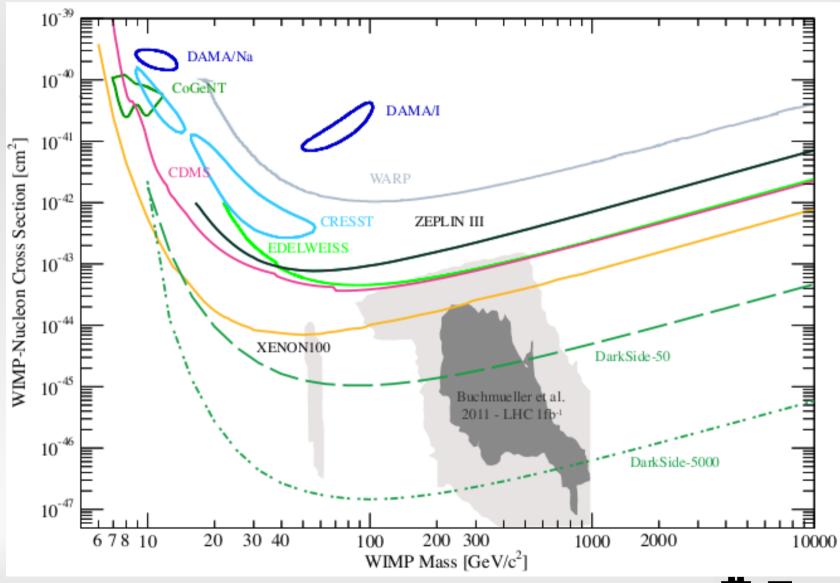
- Neutron veto tank is installed, expect first data early 2013
- Plan to run for three years
- Best estimates are for <0.5 total background events
- Vetos are already sized to hold a ~3 ton cryostat for the next generation





DarkSide-50 Predicted Sensitivity





Google image search delivers





lifeisaroad.com

